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being brought nearer to the distinst Base of the Object Glass; and an Eye less Convex, the office of a less Convex Eye-Glass: but with this difference, that the more Convex the Eye is, the easier may any part of the Object be found, and the larger and more lucid it will appear.

I have seen Saturn's Ring very plain with an Object-Glass of little more than six Foot Radius, without

an Eye-Glass.

I have also found out a way for the Presbyte to make use of an Object-Glass, by placing their Eye nearer the Lens than its Focus, by so much as their Eye is flatter than a common Eye, so as to make (as it were) the Telescope of Galileo; the flat Eye serving as a common Eye arm'd with a Concave Lens. I have so fixed the Telescope, as to make a Presbyta read at a great distance a small Print. The truth of this may be easily demonstrated, if it be requir'd.

If this Experiment be made at Sea with a very large Tube, big enough to put in the Head and move it about, and the Object Glass be also large, it may not perhaps be difficult to observe the Eclipses of the Satellites of Jupiter, which I would recommend to the Consideration of those that would try for the Longi-

tude by such like Observations.

VI. New and accurate Tables for the ready Computing of the Eclipses of the first Satellite of Jupiter, by Addition only. By the Reverend Mr James Pound, R. S. S.

N Numb. 214. of these Transactions, for the Months of Novem and Decem. 1694. we exhibited an Epitomy of Mr. Cassini's curious Tables then newly published.

lished for computing the Eclipses of the first Satellite of Jupiter, without the help of any other Numbers. The ease of this Calculus gave great satisfaction to those that delight in Telescope observations; and has been of good use to encourage Astronomers to ascertain the Geographical Longitudes of many places, by help of these Eclipses; whose frequency seems to afford us the

properest means for that purpose.

But it being now 26 Years fince those Tables were published, length of Time has discovered that this Satellites motion is a small matter swifter than M. Cassini had supposed it; and the Reverend Mr. Pound being provided with all the Qualifications requisite for such a Work, has of late apply'd himself to restify by frequent Observation what he found amis in the aforesaid Calculus; and withal has put it into another Form yet much more easy and compendious, by bringing what M. Cassini had given us in odd Numbers, to the Millesimals of a Circle, both as to Numb. I. which he calls Numb. A. being the mean Anomalie of Jupiter in fuch parts; as also to Numb. II. or our Numb. B. which is the distance of the mean place of Jupiter, from the true place of the Sun, and which with the addition of the Equation of Numb. B. gives the true angle of Commutation in the same Millesimals of a Circle. And having deducted from the Epoches the greatest Equarions both of Numb. A and B. he restores them by adding as much to the Equations themselves, by which means they all become Affirmative, so that the whole computation is performed by Addition only.

The Reader is supposed to be acquainted with the Method of M. Cassini's Calculus, which is at large explain'd in the aforesaid Transaction, Num. 214. For which reason this shorter Description may suffice at present.

(1023)

Epochæ Conjunctionum Primi Satellitis Cum Jove.

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Revolutiones Primi Satellitis Jovis in mensibus.

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24 18 40 23	6	64	6 11 38 10 15	168
26 13 8 59	6	69	8 6 6 46 16	1 / 2 1
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30 2 6 11	7	78	11 19 3 58 16	182
31 20 34 47	7	82	13 13 32 34 17	186
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2 15 3 23		87	20 15 26 58 18	204
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9 16 57 47		105	25 22 52 46 20	217
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Revolutiones Primi Satellitis Fovis in mensibus.

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Revolutiones Primi Satellitis Jovis in me Sibus.

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6 14 31 32	43	468	21 14 55 7	54	580
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10 3 28 44	44	476	25 3 52 18	55	588
11 21 57 20	45	480	26 22 20 54	55	593
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5 16 37 43	50	541			
7 11 6 19	51	545		61	659
9 5 34 55	51	549		62	663
11 0 3 31	51	554		62	672
12 18 32 7	52	558		63	677
14 13 0 43	521	562		,	- ' '

(1027)

Revolutiones Primi Satellitis Jovis in mensibus.

O&obris.	N.	Nu.	Novembris. N. Nu.
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I 7 52 54	63	681	16 8 16 29 74 799
3 2 21 30	64	686	18 2 45 5 74 804
4 20 50 6	64	690	19 21 13 40 75 808
6 15 18 41	65	695	21 15 42 16 75 813
8 9 47 17	65	699	23 10 10 52 76 817
10 4 15 53	65	704	25 4 39 28 76 822
11 22 44 29	66	708	26 23 8 4 76 827
13 17 13 5	66	713	28 17 36 40 77 831
15 11 41 41	67	717	30 12 5 16 77 836
17 6 10 17	67	721	
19 0 38 53	67	726	Decembris.
20 19 7 29	68	730	0 12 5 16 77 836
22 13 36 5	68	735	2 6 33 52 78 840
24 8 4 41	69	739	j 4 1 2 28 78 845
26 2 33 17	69	744	5 19 31 4 78 849
27 21 1 53	69	749	7 13 59 40 79 854
29 15 30 29	70	753	9 8 28 16 79 859
31 9 59 5	70	758	11 2 56 52 80 863
Novembris.	- 1		13 21 25 28 80 868
1			14 15 54 4 80 873
	70	758	16 10 22 40 81 877
		762	18 4 51 16 81 882
, , ,	3	767	19 23 19 52 82 886
		772	21 17 48 28 82 891
1 , , , , , , , , , , , ,		776 781	23 12 17 4 82 897
9 6 22 5	1		25 6 45 40 83 900
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(1028)
Prima Aquationes Conjunctionum Primi Satellitis cum Jove.

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Prima Aquationes Conjunctionum Primi Satellitis cum Jove.

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Secunda

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Secundæ Æquationes Conjunctionum Primi Satellitis cum Jove.

Addenda.

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Nu A.	1	Equa-	Num.	Nu. A.				Nu.	1		idu nes.		•		idu nes		Ти. 1 .	1		du nes.
	- -	, ,,			H	٠,	"		B	,	11		H	. ,	• • • • • • • • • • • • • • • • • • • •			H	. ,	"
) 3	30	1000	0	[[5	9	250	įI	7	0	500	I	5	9	7	50	I	7	46
20	3	29	980	10	I	4	56	260	I	7	15	510	I	4	53	7	60	r	7	57
4	3	28	960	1	r	4	44	270	1	7	3 I	520	1	4	39	7	70	I	8	7
60	1 -	25	940		I	4	33	280	1	7	45	530		4	26	11 '	80	•	8	15
80	1 -	19	920		I	4	23	1290	1	7	5 7	540		4	15		90	1	8	22
100	3	12	900	50	I	4	13	300	I	8	_7	550	1	4	7	8	00	ı	8	26
120	3	4	880		I	4	7	310	1	8	15	560	r	4	3		10		8	28
140		,		4 '	ł	4	4	320		8	22	570	1	4	1		20	•	8	30
160		•	840			4	2	330		8	27	580		4	0		३०		8	28
180	ı	٠٠.	820	90		4	0	340		8	28	590		4	3		40		8	26
200	2	22	800	100		4	2	350	-	8	29	600	I	4	_7	8	50	<u>'</u>	8	22
220	2	10	780		I	4	3	360		8	27	510	1	4	13		60		8	16
240		57	760		I	4	6	370	I	8	24	620		4	23		70		8	8
260	3	44	740	- 1	I	4	12	380	1	8	17	530		4	35		80		8	0
280	ł	30	. 1	140		4	21	390	I	8	9	640	I	4	49		90		7	50
300	I	17	700	150		4	31	40 0	I	7	58	650	1	5	4	19	0 0	I	7	37
320	I	5	680	160	1		42	410	I	7	46	660	1	5	19	19	10	I	7	22
340		53		170		4	55	420	,	7	31		1	7	36	•	20		7	8
360		41		180		5	9	430		7	14	1.	I	5	54	1	30		6	55
380		31		190		-	23	440		6	58	690		6	10	1	40		6	40
400	10	22	600	200	<u>I</u>	5	39	450	1	6	40	700	1	6	28	9	50	I	6	23
420		14			Į	5	55	460	I	6	20	710	I	6	46	9	6 0	ŧ	6	8
440		8	- (220		6	11	470		6	2	720	I	7	2		70	í	5	54
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480	ı	2	520	• 1	I		43	490		5	26	740	I	7	33	1 -	90	i	5	22
, 200	1,	0'	100	250'	7	7	0;	1500	' #	5	91	1750		7	46	110	OO.	1	5	9

The Use of the foregoing T A B L E S.

HE Eclipses of the first Satellite of Jupiter, as has been already said, assord the best means of determining the Longitude of places on the Land, where Telescopes of a convenient length may be used; thirteen of these Eclipses happening every 23 Days; but it is requisite that the Observer know near the matter when these opportunities offer themselves, least on the one hand he let them slip, or else grow weary by a too long attendance on them.

Those therefore who are curious to observe them, may readily compute the times of the Immersions or Emersions of this Eatellite, and that with great exactness, by the following very sho t Precepts, which

admit of no Exception or Caution, viz.

Out of the first Table take the Epoche for the Year. with its corresponding Numb. A and Numb. B; and to them add, out of the Tables of Months, the Day, Hour, Minute and Second, nearest less than the time of the Eclipse you seek for, together with its Num. A and B: the Sum of the times is the mean time of the middle of the Ecliple 2. With Num A thus colletted take out the first Æquation of the Conjunctions; as also the Aguation of Num. B. always to be added to Num. B. before found. 3. With Num B fo equated, take out the second Aguation of the Conjunctions; and in the last Table, the third Æquation, as also the Semi-duration of the Feliple answering to Num. A. 4. To the mean time of the middle of the Eclipse, add all those three Aquations; the 5um shall be the true equated time of the middle of the Eclipse fought. 5. If Num. B. equated be less than 500, subliract

the Semiduration, and you will have the time of the Immersion, or if it be more than 500, adding the same, it will give the time of the Emersion

But Note, the times thus found are equal time, still to be reduced to the Apparent: and that in the Biffex-tile Year, after February, one Day is to be deducted

from the Day of the Month.

The less skilful may perhaps be p'eas'd with an Example or two, which may serve them to imitate. Let it be required to find the time of the immertion of this Satellite into Jupicer's shadow, November the 9th 1719. in the Morning. The Work stands thus,

	D. h.	i	Nu. A.	Nu. B.	i
1719.			872	396	
Novemb.	7.11.	53.29	72	776	
Conj. Med.	8.18.	4.42	944	172	
Æquat. 1.		51.53		10.	Æq.B.
Æquat. II.				182	B. Æquat.
Æquat.III.		3.26		-02	De zuquate
	8.19.	10.27			
	1.	6.33	Semidur. Sa	bst.	
Novemb.	8.18.	3 • 5 4			

So that by this Calculus, on the ninth of Novembat 4 Minutes after 6 in the Morning, equal Time, may be seen the Immersion of this Satellite into Jupiter's shadow.

Another Example shall be of the Emersion on the fifth of April 1720, viz.

			<i>-</i>		
	D. h.	,	Nu.	A. Nu. I	3.
	0.20.			6 310	
April	4.13.	44.2	2 Bils. 2	2 244	
Conj. Med.	5.10.	07.0	97		
Æquat. I.		44 . 1		1.3	Æq. B.
Æquat. II.		0.45	5	567	B. Æquat.
Æquat. III.		3.29			D. andage
	ı.	5 · 40	Semidur	. Add.	
April	5.12.	01.09)		

Hence it appears that at one Minute after Midnight following the fifth of April, equal Time, will happen the Emersion required. Nor do we doubt but that the Event will very nearly answer.

Lastly, it may not be amisshere to inform the Reader, that we have learnt, by the experience of many Years Observation, that the second inequality of this Satellite proceeds from the progressive Propagation of Light, and is common to all the rest of the Satellites: Light, being found to proceed in about seven Minutes of time as far as from the Sun to the Earth, whether with an equable motion or otherwise is still a question. For this reason we have added a Third Aquation, whereby to account for the greater distance of Jupiter from the Earth in Aphelio than in Perihelio, as the Second Aquation answers to the greater distance of the Planet when near the Conjunction of the Sun, than when near his Opposition.

F I N I S.

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